Sparse Endmembers and Demixing

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Figure: Martin Ehler

Outline

- Hyperspectral data and endmembers
 - Hyperspectral data
 - Endmembers
- Sparse endmembers
 - Models
 - Theoretical underpinnings
 - Selecting the endmembers
- Results
 - Urban
 - Smith
 - Final remarks

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Color image







Red



Hyperspectral imagery data

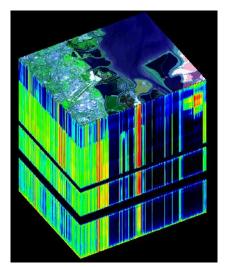


Hyperspectral camera in action



Figure: http://www.diamond-sensing.com/uploads/media/Hyperspectral.jpg

Hyperspectral data cube



Overview of hyperspectral imagery data

- Hyperspectral imagery (HSI) data is characterized by the narrowness and contiguous nature of the measurements.
- HSI data sets are spectrally overdetermined, and thus provide ample spectral information to distinguish between spectrally similar (but unique) materials.
- HSI data sets can be useful for the following purposes:
 - target detection
 - material classification
 - material identification
 - mapping details of surface properties

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Notations

- Assume our HSI data set is an $n_1 \times n_2 \times d$ cube.
 - n_1 , n_2 spatial dimensions.
 - $n = n_1 n_2 =$ number of pixels.
 - d is the spectral dimension (so d wavelengths measured).
- d is usually large, e.g., d > 100.
- n is usually very large, e.g., $n = \mathcal{O}(10^5)$ or even $n = \mathcal{O}(10^6)$.
- Let $\mathcal{X} = \{x_i\}_{i=1}^n \subset \mathbb{R}^d$ denote the pixel vectors of the HSI data cube in set form.
- Let $X = [x_1 \ x_2 \ \cdots \ x_n]$ be a $d \times n$ matrix where the columns x_i of X are the pixel vectors of the HSI data cube.

Endmembers

Definition

Endmembers are a collection of a scene's constituent spectra. If $\mathcal{E} = \{e_i\}_{i=1}^s \subset \mathbb{R}^d$ are endmembers corresponding to a data set \mathcal{X} , then there is some representation of each $x_i \in \mathcal{X}$ in terms of the elements of \mathcal{E} .

- Let $E = [e_1 \ e_2 \ \cdots \ e_s]$ be a $d \times s$ matrix where the columns e_i of E are the endmembers.
- s is usually small, e.g., s < d.
- Many algorithms find the endmembers from within the data, so that $\mathcal{E} \subset \mathcal{X}$.
- One alternative is to find endmembers from a spectral library, L, that can be used for multiple data sets.



Linear mixture model

• Given a data set $\mathcal X$ and corresponding endmembers $\mathcal E$, the linear mixture model states that:

$$x_i = \sum_{j=1}^s \alpha_{i,j} e_j + z_i, \quad ext{for all } x_i \in \mathcal{X}.$$

- $\alpha_{i,j} \geq 0$ for all $i = 1, \ldots, n$ and $j = 1, \ldots, s$.
- $\sum_{i=1}^{s} \alpha_{i,j} = 1$ for all i = 1, ..., n.
- $z_i \in \mathbb{R}^d$ is a noise vector.

Visualization of the linear mixture model

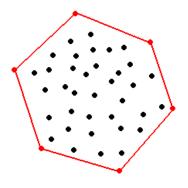


Figure: The linear mixture model



Examples of endmember algorithms

Some endmember algorithms are the following:

- N-FINDR [M. Winter]: Finds the simplex of maximal volume that contains the data set X; the vertices of this simplex are the endmembers.
- SVDD [D. Tax and R. Duin]: Obtains a spherical shaped boundary around the data set X; support vectors, or endmembers, are derived from this description.
- Pixel Purity Index [J. Boardman]: Repeatedly projects
 d-dimensional scatter plots onto a random unit vector; the
 extreme pixels in each projection are recorded and the total
 number of times each pixel is marked as extreme is noted.

Endmember coefficients

- After one finds an endmember set \mathcal{E} , the coefficients $\{\alpha_{i,j}\}_{i,j=1}^{n,s}$ must be computed.
- Two common ways of computing the coefficients are the following:
- Minimum error:

$$\alpha_{i,\cdot} = \arg\min_{\tilde{\alpha}} \|x_i - E\tilde{\alpha}\|_{\ell^2}$$

② Sparse: let $\tau_i > 0$,

$$\alpha_{i,\cdot} = \arg\min_{\tilde{\alpha}} \|x_i - E\tilde{\alpha}\|_{\ell^2}^2 + \tau_i \|\tilde{\alpha}\|_{\ell_1}$$

• Note when solving either minimization problem, $\tilde{\alpha}$ is subject to the constraints of the linear mixture model.



A look ahead

- Even if one uses the sparse coefficient model, the endmember algorithm itself does not necessarily select the endmembers with sparsity in mind!
- The endmember algorithm presented in the next section is based on the sparse coefficient model.
- We will be searching for endmembers as a subset of \mathcal{X} , i.e., we assume that $\mathcal{E} \subset \mathcal{X}$.

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A simplistic model

- Assume the linear mixture model, and furthermore suppose that $z_i = 0$ for all $i = 1, \ldots, n$.
- Define an $n \times n$ weight matrix, $W = (w_{i,j})$, as follows.
- Let $c: \mathbb{R}^n \longrightarrow \mathbb{R}$ be a cost function.
- If possible.

$$w_{i,\cdot} = \arg\min_{\tilde{w}} c(\tilde{w}), \text{ subject to:}$$

- $\sum_{j=1}^{n} \tilde{w}_{j} x_{j} = x_{i}$ $\tilde{w}_{j} \geq 0$ for all $j = 1, \dots, n$.
- $\sum_{i=1}^{n} \tilde{w}_{i} = 1$
- $\tilde{w}_i = 0$
- Otherwise, $w_i = \delta_i$.
- We can extract the endmembers from the rows of W. Namely x_i is an endmember if its corresponding weight row satisfies $w_{i,\cdot} = \delta_i$.



Observations on the previous model

• Notice that in the previous model we are representing each $x_i \in \mathcal{X}$ in terms of the dictionary, or finite frame,

$$\mathcal{X}^{(i)} = \mathcal{X} \setminus \{x_i\},\,$$

subject to the constraints of the linear mixture model.

• If the endmember set $\mathcal{E} \subset \mathcal{X}$ is sparse enough in the dictionary \mathcal{X} (and thus in each $\mathcal{X}^{(i)}$ as well), then we could set the cost function as

$$c(\tilde{w}) = \|\tilde{w}\|_{\ell^0} = |\operatorname{supp}(\tilde{w})|,$$

and expect that the support of each $w_{i,\cdot}$ lies within \mathcal{E} .

• Thus we could extract the endmembers from the columns of the weight matrix W as well! In particular, if $\operatorname{supp}(w_{\cdot,i}) \neq \emptyset$, then $x_i \in \mathcal{E}$.



Another simplistic model

- In reality, the endmembers will not be guite so apparent.
- Assume only part of the linear mixture model: remove the convexity (i.e. the sum to one) constraint.
- Once again assume that $z_i = 0$ for all i = 1, ..., n.
- Suppose that $s < d \ll n$, which makes \mathcal{E} sparse in \mathcal{X} .
- Define an $n \times n$ weight matrix $W = (w_{i,i})$ as follows:

$$w_{i,\cdot} = \arg\min_{\tilde{w}} \|\tilde{w}\|_{\ell^0}$$
, subject to:

- $\tilde{w}_i = 0$



Observations on the second model

- Notice that even the endmembers $e_i \in \mathcal{E}$ will have such a representation in the dictionary $\mathcal{X}\setminus\{e_i\}$.
- However, this representation will be a misrepresentation!
- For each $x_i \notin \mathcal{E}$ though, the support of $w_{i,\cdot}$ will be contained in \mathcal{E} .
- Thus if the weight of the 'good' representations outweighs the 'bad' representations, then we will still extract the endmembers \mathcal{E} from the columns of W.
- In particular, we know that for each $x_i \notin \mathcal{E}$, we have $\|w_{\cdot,i}\|_{\ell^0} \leq s$.
- For $x_i \in \mathcal{E}$ though, we know that $||w_{\cdot,i}||_{\ell^0} \leq n s$.
- Also, due to the fact that $s \ll n$, we almost certainly have $\|w_{\cdot,i}\|_{\ell^0} > s$.
- Thus we could extract the endmembers by selecting the $x_i \in \mathcal{X}$ corresponding to the largest $||w_{\cdot,i}||_{\ell^0}$.



Synthetic weight matrix

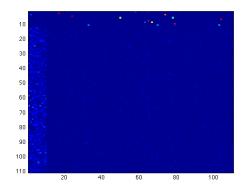


Figure: Weight matrix of synthetic data set



Adding back in the noise

- In practice of course, $z_i \neq 0$.
- Furthermore, the ℓ^0 pseudo–norm is computationally intensive, and so we turn to the ℓ^1 norm.
- In order to account for these two issues, we define our $n \times n$ weight matrix $W = (w_{i,j})$ as follows:

$$w_{i,\cdot} = \arg\min_{\tilde{w}} \|\tilde{w}\|_{\ell^1}, \text{ subject to:}$$
 (1)

- $\bullet \|x_i X\tilde{w}\|_{\ell^2} \leq \delta_i,$
- $\tilde{w}_j \geq 0$ for all $j = 1, \ldots, n$,
- $\bullet \ \tilde{w}_i = 0.$
- Note that (1) can be replaced with:

$$w_{i,\cdot} = \arg\min_{\tilde{w}} \|x_i - X\tilde{w}\|_{\ell^2}^2 + \lambda_i \|\tilde{w}\|_{\ell^1}, \text{ subject to:}$$
 (2)

- $\tilde{w}_j \geq 0$ for all $j = 1, \ldots, n$,
- $\bullet \ \tilde{w}_i = 0.$



Comments on sparse endmember extraction

 We use the second formulation of the noisy minimization, usually setting

$$\lambda_i = (.01) \cdot (X^{(i)})^t x_i.$$

- λ_i controls the density of the weight matrix. Large values of λ_i will give less vectors in \mathcal{X} as possible endmembers.
- In practice, if the non-negativity constraint is not enforced, the percentage of non-negative weights is around 0.01%. Therefore this constraint can usually be removed in order to speed up run time.
- There are (at least) two questions:
- Are there any theoretical underpinnings to this approach?
- Which columns of W are the most significant?



Deterministic results

- Let Φ be a $d \times n$ dictionary.
- Let $x_0 \in \mathbb{R}^d$ be a signal that has a sparse representation $\alpha_0 \in \mathbb{R}^n$ in Φ , i.e. $x_0 = \Phi \alpha_0$.
- Suppose all we have observed though is $x = x_0 + z$, where $z \in \mathbb{R}^d$ is noise vector satisfying $||z||_{\ell^2} \le \varepsilon$.
- Define $\hat{\alpha}_{\delta,\varepsilon}$ as

$$\hat{\alpha}_{\delta,\varepsilon} = \arg\min_{\tilde{\alpha}} \|\tilde{\alpha}\|_{\ell^1} \ \text{ subject to } \ \|x - \Phi\tilde{\alpha}\|_{\ell^2} \leq \delta.$$

Deterministic results continued

Theorem (Donoho, Elad, Temlyakov)

If Φ and α_0 satisfy certain sparsity conditions, then

$$\|\hat{\alpha}_{\delta,\varepsilon} - \alpha_0\|_{\ell^2} \le C \cdot (\varepsilon + \delta).$$

Theorem (Donoho, Elad, Temlyakov)

If we exaggerate the noise level by setting $\delta = C' \cdot \varepsilon$, where C' is a particular constant dependent on Φ and α_0 , then

$$\operatorname{supp}(\hat{\alpha}_{\delta,\varepsilon}) \subset \operatorname{supp}(\alpha_0).$$

Probabilistic results

- Again let Φ be a d × n dictionary.
- Let $x \in \mathbb{R}^d$ be our observed signal such that $x \approx \Phi \alpha$.
- Define $\hat{\alpha}_{\varepsilon}$ as

$$\hat{\alpha}_{\varepsilon} = \arg\min_{\tilde{\alpha}} \|\tilde{\alpha}\|_{\ell^{1}} \ \ \text{subject to} \ \ \|x - \Phi\tilde{\alpha}\|_{\ell^{2}} \leq \varepsilon.$$

Theorem (Donoho)

There exists $\rho > 0$ and C > 0 so that for all large d, the overwhelming majority of all $d \times n$ matrices Φ have the following property: For each vector x admitting an approximation $||x - \Phi \alpha_0||_{\ell^2} \le \varepsilon$, by some vector α_0 obeying $\|\alpha_0\|_{\ell^0} < \rho d$, then

$$\|\hat{\alpha}_{\varepsilon} - \alpha_0\|_{\ell^2} \le C \cdot \varepsilon.$$



Endmember selection

Given the weight matrix W, we select the endmembers $\mathcal{E} \subset \mathcal{X}$ according to two criterion on the columns of W.

- Support size should be large.
- Intensity per weight should also be large.

The exact method of selection

We rank the columns of *W* according to the two criterion.

- First sort the columns of W according to their ℓ^0 pseudo-norm, $\|w_{\cdot,i}\|_{\ell^0}$. The larger the support, the better the rank (i.e., the more important that column is).
 - Columns with empty support are automatically discarded at this step.
- ② Similarly, sort the columns according to the value of $\|w_{\cdot,i}\|_{\ell^2}/\|w_{\cdot,i}\|_{\ell^0}$. The larger the intensity per weight, the better the rank in this ordering.
- **3** Combine the two rankings to form a final ordering on the columns of W. The s highest ranked columns in this ordering correspond to the s endmembers in \mathcal{E} .

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Small subset of Urban

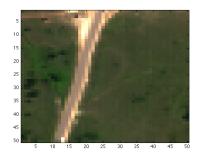


Figure: Small subset of Urban

- 50×50 pixels.
- 161 spectral dimensions.



Weight matrix of the Urban subset

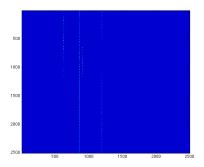
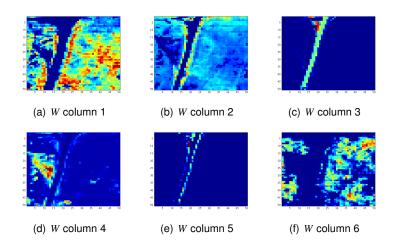


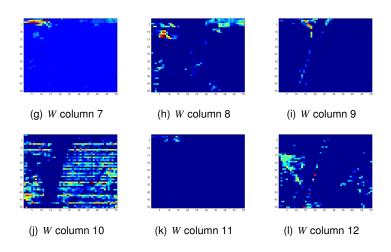
Figure: Weight matrix of the Urban subset

- Number of nonzero columns: 31.
- Number of columns $w_{\cdot,i}$ such that $||w_{\cdot,i}||_{\ell^0} \ge 10$: 18.

Weight columns of the Urban subset



Weight columns of the Urban subset



Endmembers of the Urban subset

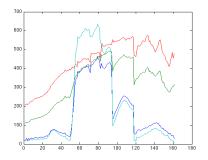
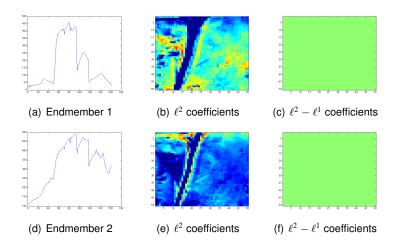
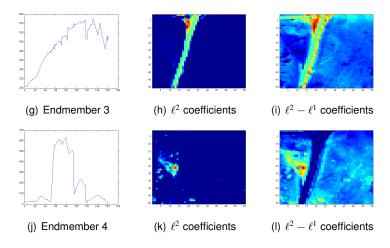


Figure: Endmembers of the Urban subset

Coefficients of the Urban subset



Coefficients of the Urban subset



Sparsity statistics for the Urban subset

The average ℓ^0 norm per pixel of each coefficient set:

- ℓ^2 coefficients: 1.9044.
- $\ell^2 \ell^1$ coefficients: 1.9044.

Note though that the ℓ^2 coefficients though do a better job of putting different materials with different endmembers!

Large data sets

- For large data sets, e.g. $n \ge 10^4$, computing the weight matrix W may be too time intensive.
- In order to get around this problem, we sample the data set uniformly at random; call this sample $\mathcal{Y} \subset \mathcal{X}$.
- We then compute the weight matrix for \mathcal{Y} , and in turn select the endmembers from \mathcal{Y} as well.
- \bullet The coefficients for the whole data set ${\mathcal X}$ are then computed from these endmembers.

Urban

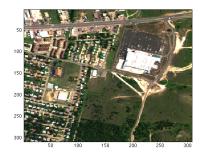


Figure: Urban: http://www.agc.army.mil/Hypercube/index.html

- 307 × 307 pixels.
- 161 spectral dimensions.
- Sample size: 4000 pixels



Weight matrix of Urban

The weight matrix, *W*, of Urban had the following statistics:

- Number of nonzero columns: 90.
- Number of columns $w_{\cdot,i}$ such that $||w_{\cdot,i}||_{\ell^0} \ge 10$: 50.

Endmembers of Urban

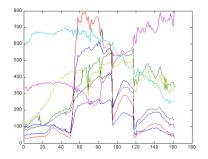
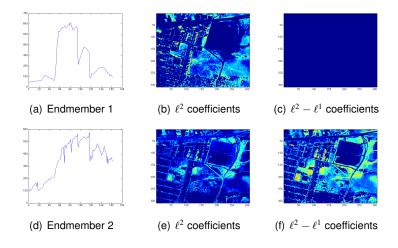
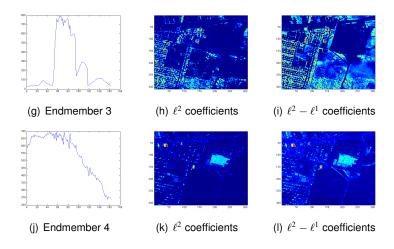


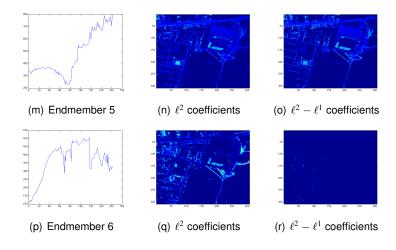
Figure: Endmembers of Urban

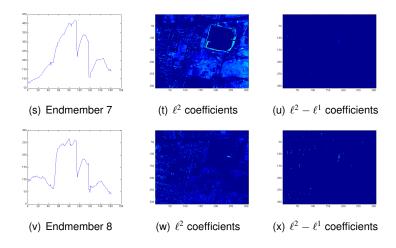
Urban



Urban







Sparsity statistics for Urban

The average ℓ^0 norm per pixel of each coefficient set:

- ℓ^2 coefficients: 3.8503.
- $\ell^2 \ell^1$ coefficients: 2.7217.

Smith



Figure: Smith

- 679×944 pixels (497182 nonzero pixels).
- 110 spectral dimensions.
- Sample size: 5000 pixels.



Weight matrix of Smith

The weight matrix, W, of Smith had the following statistics:

- Number of nonzero columns: 23.
- Number of columns $w_{\cdot,i}$ such that $||w_{\cdot,i}||_{\ell^0} \ge 10$: 16.

Smith

Endmembers of Smith

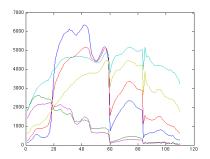
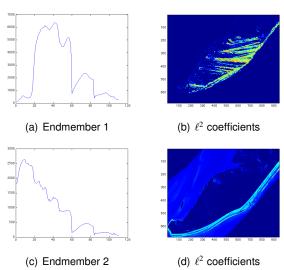
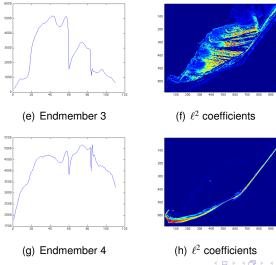


Figure: Endmembers of Smith

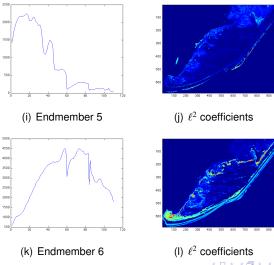
Coefficients of Smith



Coefficients of Smith



Coefficients of Smith



Sparsity statistics for Smith

• The average ℓ^0 norm per pixel of the ℓ^2 coefficients: 2.5063.

Extension to spectral libraries

- We can easily extend this method to search for endmembers from a spectral library, L.
- In fact, by computing the weight matrix, W, with dictionary

$$\Phi = \mathcal{X}^{(i)} \cup \mathcal{L},$$

we can search simultaneously for endmembers from both the given data set and the spectral library.

• Note that we would only compute weights for $x_i \in \mathcal{X}$.

For the future

Things we are working on:

- Developing more intricate and realistic models in which it is possible to obtain provable results.
- Consider smarter sampling methods that ideally would:
 - Reduce randomness in the endmember output.
 - Facilitate further gains in sparsity when computing the weight matrix, W.
- Continue to run trials on both real and synthetic data sets, and in particular, branch out to biomedical imaging.
- Systematically compare with other endmember methods.



Thank you!

Thank you for your time!